

RESOLUTION ANALYSIS BY RANDOM PROBING

Andreas Fichtner

ETH Zurich

Tristan van Leeuwen

Utrecht University

"Solving an inverse problem means to describe the infinite-dimensional space of data-fitting models."

George Backus & Freeman Gilbert, 1968

1. Why resolution analysis is becoming more and more difficult

A simple example

THE RESOLUTION MATRIX

$$R \quad m_{\text{true}} = m_{\text{est}}$$

Resolution matrix How
the true Earth is
smeared into an image.
Dimension $N \times N$.

True Earth model
Dimension N .

Estimated Earth model
Dimension N .

THE RESOLUTION MATRIX

$$\mathbf{R} \mathbf{m}_{\text{true}} = \mathbf{m}_{\text{est}}$$

Resolution matrix How
the true Earth is
smeared into an image.
Dimension $N \times N$.

True Earth model
Dimension N .

Estimated Earth model
Dimension N .

- In the days of Backus & Gilbert: $N = O(10^2) \rightarrow \mathbf{R}$ is $O(10^2)$ times larger than \mathbf{m} .

THE RESOLUTION MATRIX

$$\mathbf{R} \mathbf{m}_{\text{true}} = \mathbf{m}_{\text{est}}$$

Resolution matrix How
the true Earth is
smeared into an image.
Dimension $N \times N$.

True Earth model
Dimension N .

Estimated Earth model
Dimension N .

- In the days of Backus & Gilbert: $N = O(10^2) \rightarrow \mathbf{R}$ is $O(10^2)$ times larger than \mathbf{m} .
- Today: $N = O(10^7) \rightarrow \mathbf{R}$ is $O(10^7)$ times larger than \mathbf{m} .

THE RESOLUTION MATRIX

$$\mathbf{R} \mathbf{m}_{\text{true}} = \mathbf{m}_{\text{est}}$$

Resolution matrix How
the true Earth is
smeared into an image.
Dimension $N \times N$.

True Earth model
Dimension N .

Estimated Earth model
Dimension N .

- In the days of Backus & Gilbert: $N = O(10^2) \rightarrow \mathbf{R}$ is $O(10^2)$ times larger than \mathbf{m} .
- Today: $N = O(10^7) \rightarrow \mathbf{R}$ is $O(10^7)$ times larger than \mathbf{m} .
- As data volumes and computing power grow:
 - We can construct bigger and bigger models \mathbf{m}_{est} .

THE RESOLUTION MATRIX

$$\mathbf{R} \mathbf{m}_{\text{true}} = \mathbf{m}_{\text{est}}$$

Resolution matrix How
the true Earth is
smeared into an image.
Dimension $N \times N$.

True Earth model
Dimension N .

Estimated Earth model
Dimension N .

The problem:

- In the days of Backus & Gilbert: $N = O(10^2) \rightarrow \mathbf{R}$ is $O(10^2)$ times larger than \mathbf{m} .
- Today: $N = O(10^7) \rightarrow \mathbf{R}$ is $O(10^7)$ times larger than \mathbf{m} .
- As data volumes and computing power grow:
 - We can construct bigger and bigger models \mathbf{m}_{est} .
 - We lose our ability to quantify the quality of \mathbf{m}_{est} .

We need **scalable** methods to infer useful **aspects** of resolution.

We need **scalable** methods to infer useful **aspects** of resolution.

Objectives of this Webinar:

- Describe 2 methods to quantify resolution when **R** is too expensive to compute and too big to store.
- One method for linear problems, and one for (mildly) nonlinear problems.
- Both based on random probing techniques.

2. Estimating the number of resolved parameters

tr R

ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

m_i

- random test model vector
- Expectation: $E[m_i]=0$
- Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]

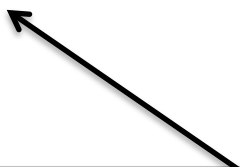
ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

m_i

- random test model vector
- Expectation: $E[m_i]=0$
- Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]

R_{ij}

- A resolution matrix



Too large to computer.
Too large to store.
Too large to comprehend fully.

ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

m_i

- random test model vector
- Expectation: $E[m_i]=0$
- Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]

R_{ij}

- A resolution matrix

$E[m_i R_{ij} m_j]$

ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

m_i

- random test model vector
- Expectation: $E[m_i]=0$
- Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]

R_{ij}

- A resolution matrix

$$E[m_i R_{ij} m_j] = R_{ij} E[m_i m_j]$$

ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

m_i

- random test model vector
- Expectation: $E[m_i]=0$
- Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]

R_{ij}

- A resolution matrix

$$\begin{aligned} E[m_i R_{ij} m_j] &= R_{ij} E[m_i m_j] \\ &= R_{ij} (E[m_i] E[m_j] + \text{cov}(m_i, m_j)) \end{aligned}$$

ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

m_i

- random test model vector
- Expectation: $E[m_i]=0$
- Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]

R_{ij}

- A resolution matrix

$$\begin{aligned} E[m_i R_{ij} m_j] &= R_{ij} E[m_i m_j] \\ &= R_{ij} (E[m_i] E[m_j] + \text{cov}(m_i, m_j)) \\ &= R_{ij} \delta_{ij} = R_{ii} = \text{tr } \mathbf{R} \end{aligned}$$

ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

m_i

- random test model vector
- Expectation: $E[m_i]=0$
- Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]

R_{ij}

- A resolution matrix

$$\begin{aligned} E[m_i R_{ij} m_j] &= R_{ij} E[m_i m_j] \\ &= R_{ij} (E[m_i]E[m_j] + \text{cov}(m_i, m_j)) \\ &= R_{ij} \delta_{ij} = R_{ii} = \text{tr } \mathbf{R} = \text{number of resolved model parameters} \end{aligned}$$

Hutchinson's method [Hutchinson, 1990]

ESTIMATING THE NUMBER OF RESOLVED PARAMETERS

- m_i
- random test model vector
 - Expectation: $E[m_i]=0$
 - Covariance: $\text{cov}(m_i, m_j)=\delta_{ij}$ [uncorrelated components]
- R_{ij}
- A resolution matrix

$$\begin{aligned} E[m_i R_{ij} m_j] &= R_{ij} E[m_i m_j] \\ &= R_{ij} (E[m_i] E[m_j] + \text{cov}(m_i, m_j)) \\ &= R_{ij} \delta_{ij} = R_{ii} = \text{tr } \mathbf{R} = \text{number of resolved model parameters} \end{aligned}$$

Very simple recipe:

- Choose a random test model \mathbf{m} .
- Try to recover this model in a synthetic inversion [i.e. compute $\mathbf{m}_{\text{est}} = \mathbf{R}\mathbf{m}$].
- Multiply the result with \mathbf{m} itself: $\mathbf{m}^T \mathbf{m}_{\text{est}} = \mathbf{m}^T \mathbf{R}\mathbf{m}$.
- Average over some random realisations.
- **The resolution matrix itself never has to be computed!**

OTHER RANDOM PROBING TECHNIQUES

Hutchinson, M. F. (1990), A stochastic estimator of the trace of the influence matrix for Laplacian smoothing splines, *Comm. Stat. Sim.*, 19, 433–450.

An, M. (2012), A simple method for determining the spatial resolution of a general inverse problem, *Geophys. J. Int.*, 191, 849–864.

Avron, H., and S. Toledo (2011), Randomized algorithms for estimating the trace of an implicit symmetric positive semi-definite matrix, *J. Ass. Comp. Mach.*, 58, doi:10.1145/1944,345.

Drineas, P., R. Kannan, and M. W. Mahoney (2006), Fast Monte Carlo algorithms for matrices II: Computing a low-rank approximation to a matrix, *SIAM J. Comput.*, 36, 158–183.

Frieze, A., R. Kannan, and S. Vempala (2004), Fast Monte Carlo algorithms for finding low-rank approximations, *J. Assoc. Comput. Mach.*, 51, 1025–1041.

Halko, N., P. G. Martinsson, and J. A. Tropp (2011), Finding structure with randomness: Probabilistic algorithms for constructing approximate matrix decompositions, *SIAM Review*, 53, 217–288.

MacCarthy, J. K., B. Borchers, and R. C. Aster (2011), Efficient stochastic estimation of the model resolution matrix diagonal and generalized cross-validation for large geophysical inverse problems, *J. Geophys. Res.*, 116, doi: 1029/2011JB008,234.


Trampert, J., and A. Fichtner (2013), Resolution tests revisited: The power of random numbers, *Geophys. J. Int.*, 192, 676–680.

3. Random probing for resolution analysis in tomography

Estimating position- and direction-dependent resolution lengths.

POINT-SPREAD FUNCTIONS

- Misfit χ in the vicinity of the optimal model \mathbf{m} :

$$\chi(\mathbf{m} + \delta\mathbf{m}) = \chi(\mathbf{m}) + \frac{1}{2}\delta\mathbf{m}^T \mathbf{H}(\mathbf{m}) \delta\mathbf{m}$$


Hessian operator

Inverse posterior covariance [assuming Gaussian errors]

Column: point-spread function

H is too expensive to compute and store.

- **But we can** infer properties of **H** from its application to random test models.

RANDOM PROBING PRINCIPLE

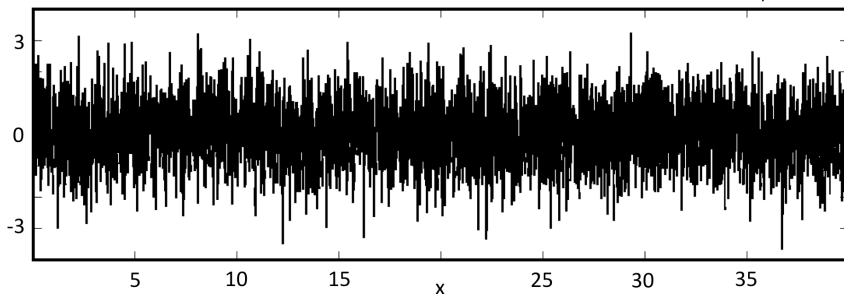
- Assume **H** is Gaussian [for simplicity and illustration]:

$$H(x; y) = \frac{1}{\sqrt{2\pi} \sigma} e^{-\frac{1}{2\sigma^2} (x-y)^2}$$

$$h(y) = \int H(x; y) v(x) dx$$

random test model

random test model $v(x)$



RANDOM PROBING PRINCIPLE

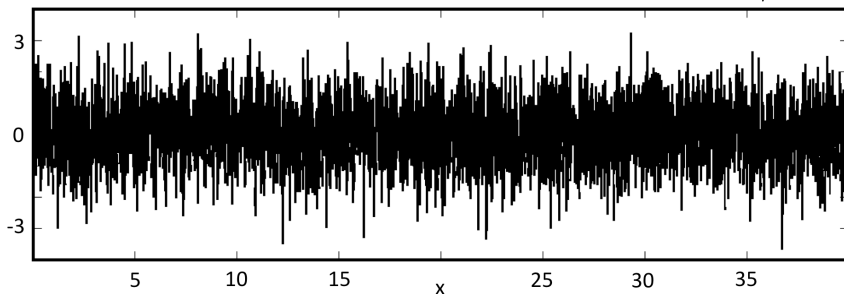
- Assume \mathbf{H} is Gaussian [for simplicity and illustration]:

$$H(x; y) = \frac{1}{\sqrt{2\pi} \sigma} e^{-\frac{1}{2\sigma^2} (x-y)^2}$$

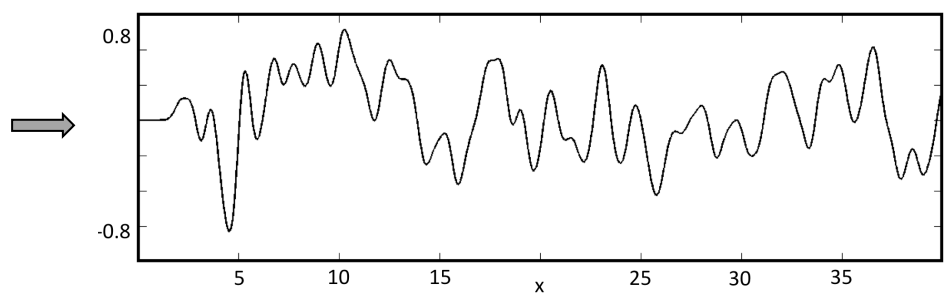
$$h(y) = \int H(x; y) v(x) dx$$

random test model

random test model $v(x)$



smoothed version of $v(x)$ [$h=H \cdot v$]

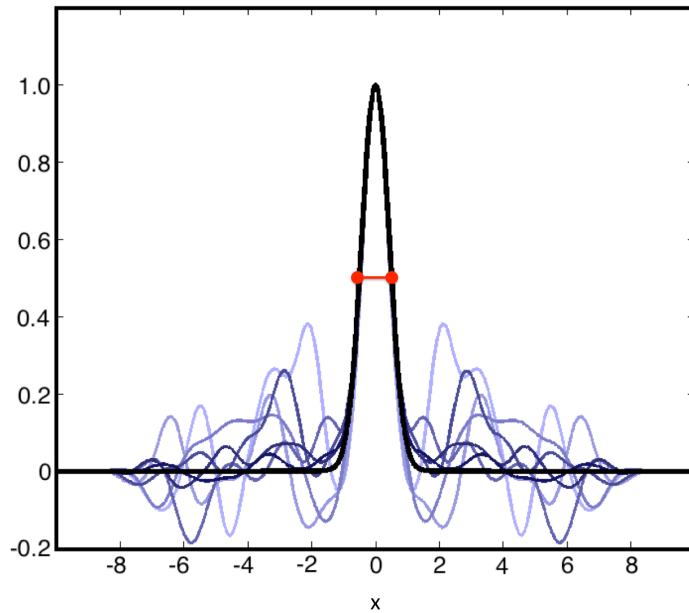


- Length scales of \mathbf{h} contain information on length scales of \mathbf{H} .

RANDOM PROBING PRINCIPLE

- **Auto-correlation** of the output **h** [averaged over many realisations]:

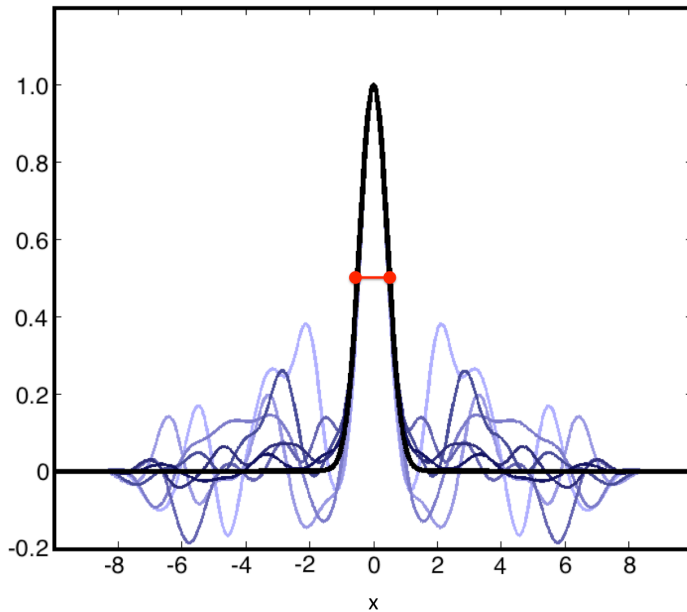
average auto-correlations of h
[for 1, 2, 3, 5, 10, 20, 50 samples]



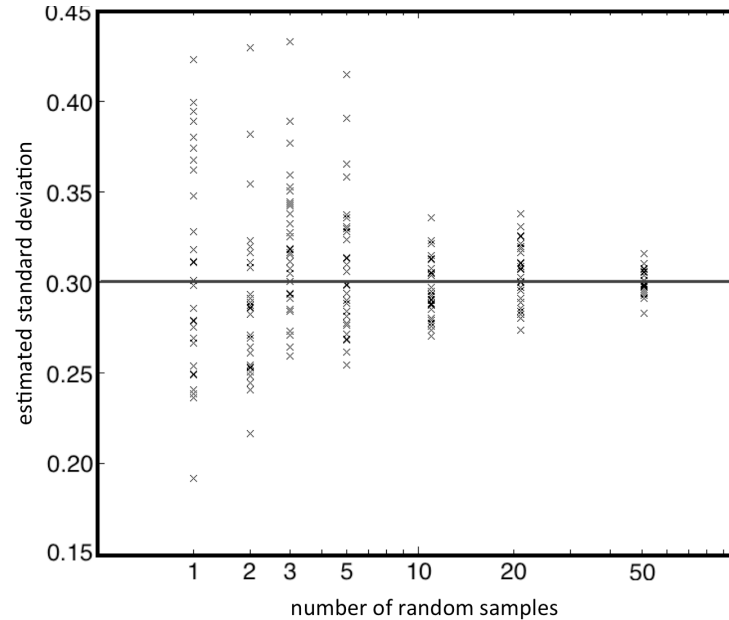
RANDOM PROBING PRINCIPLE

- **Auto-correlation** of the output **h** [averaged over many realisations]:
- **Asymptotically**: width of auto-correlation = $\sqrt{2}$ • width of **H**

average auto-correlations of **h**
[for 1, 2, 3, 5, 10, 20, 50 samples]



estimated with of **H**
[for 1, 2, 3, 5, 10, 20, 50 samples]



PRELIMINARY CONCLUSIONS

1. Resolution and correlations

- The Hessian acts as a smoother of random functions.
- The smoothed functions carry information on **resolution**.
- Can be **extracted with correlations**.

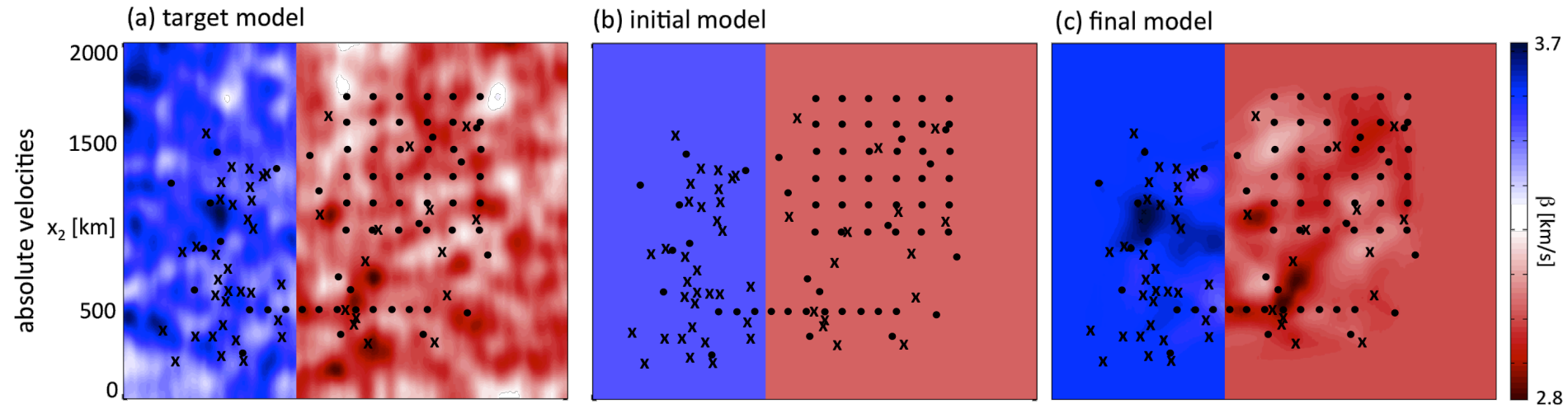
2. Convergence

- Correlations themselves may require large sample sizes to converge.
- The **width** of the correlation **converges extremely quickly**.
- Useful **resolution proxies** may already be obtained with **very few samples**.

Synthetic full-waveform inversion in 2D

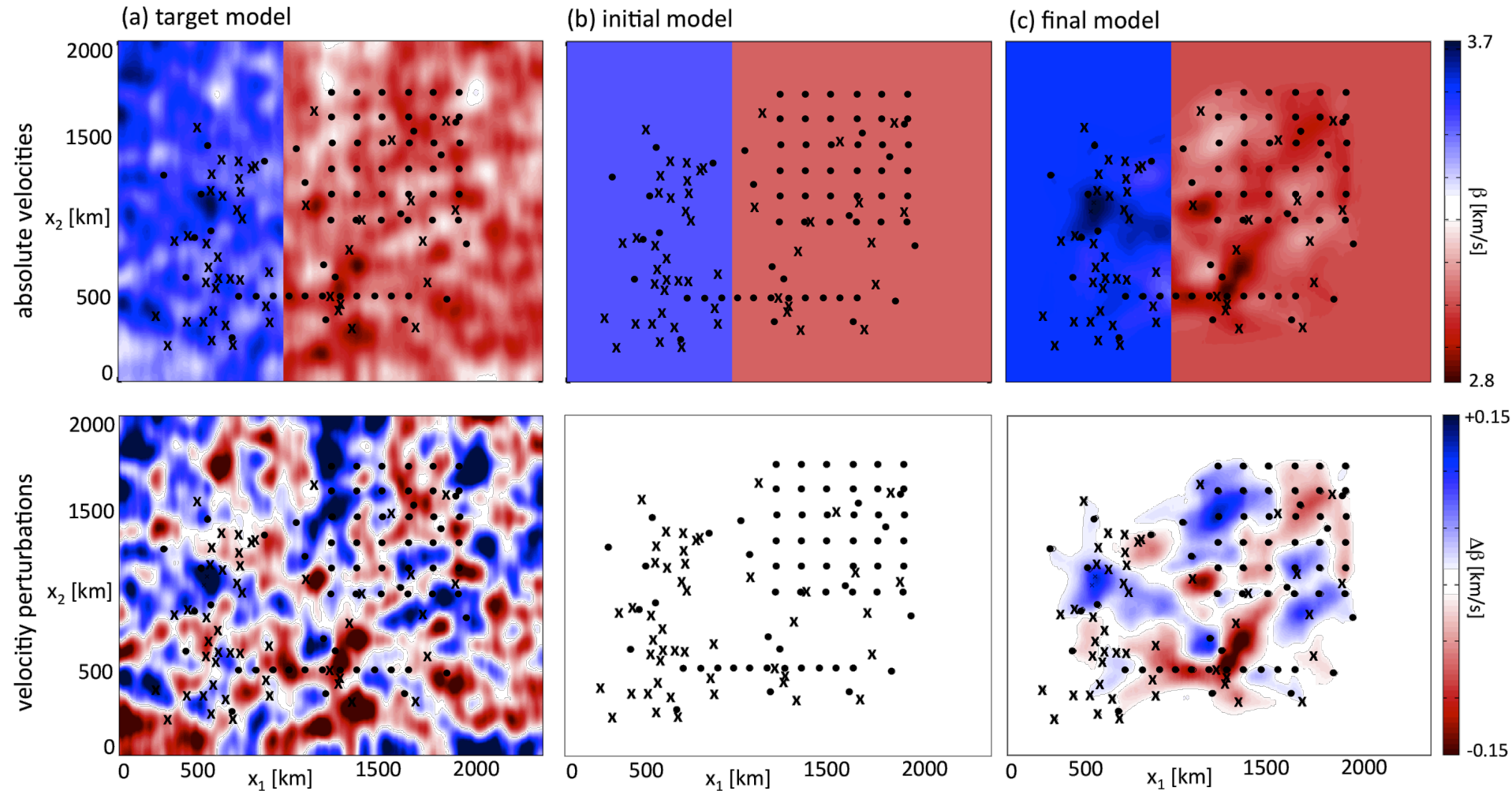
SYNTHETIC EXAMPLE IN 2D

- Synthetic inversion setup



SYNTHETIC EXAMPLE IN 2D

- Synthetic inversion setup

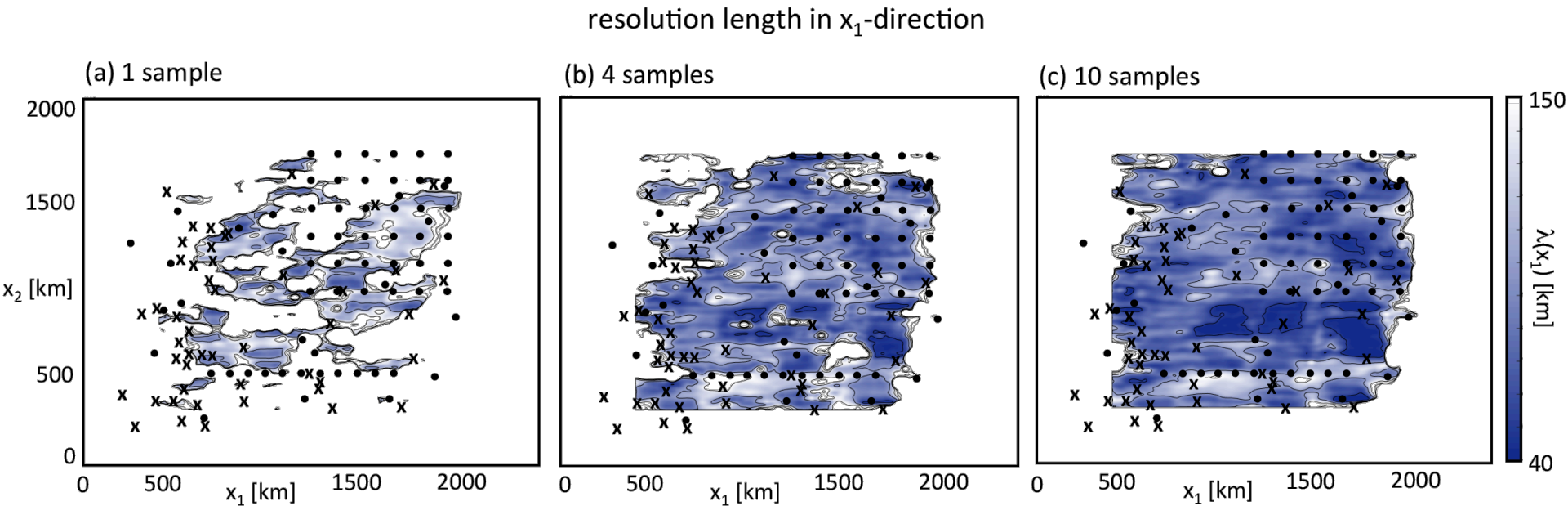


SYNTHETIC EXAMPLE IN 2D

- Application of random test models to the Hessian via **second-order adjoints**
- Local auto-correlation of the output in different directions.

SYNTHETIC EXAMPLE IN 2D

- Application of random test models to the Hessian via second-order adjoints.
- Local auto-correlation of the output in different directions.
- Estimated width of the point-spread functions in **x_1 -direction** [resolution length].

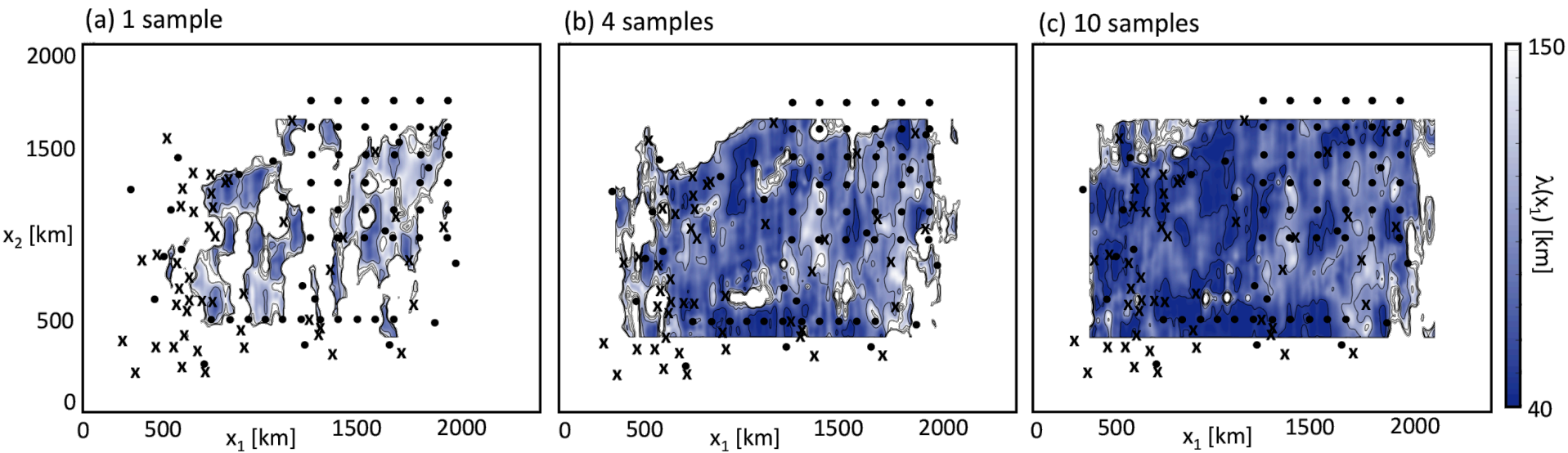


- Around 5-10 samples to converge.
- Resolution is strongly heterogeneous.

SYNTHETIC EXAMPLE IN 2D

- Estimated width of the point-spread functions in **x_2 -direction** [resolution length].

resolution length in x_2 -direction



Real-data application

INVERSION SETUP

Technical summary:

Data

- 52 earthquakes, >1000 stations
- body waves, surface waves, ...
- periods: **10 – 150 s**

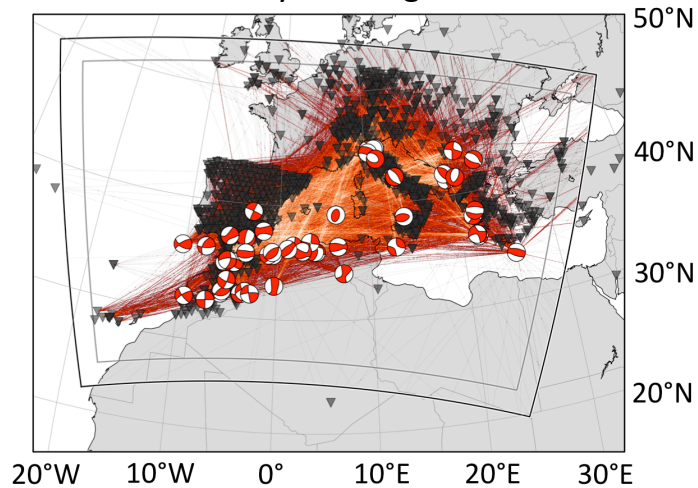
Forward modelling

- spectral elements
- **3D visco-elastic, anisotropic**

Inversion

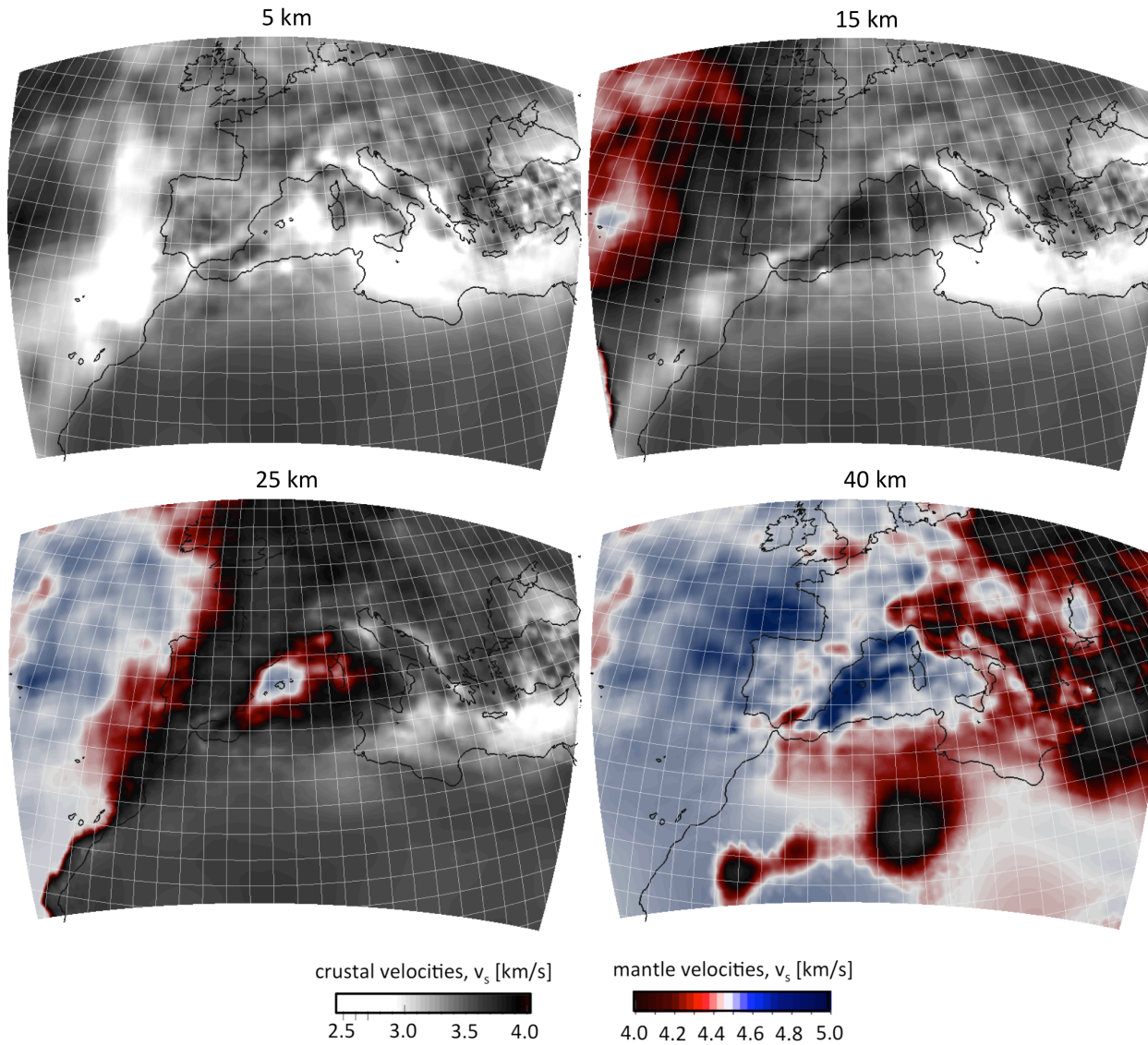
- initial model from previous European FWI
- adjoint-based CG
- invert for v_{sh} , v_{sv} , v_p , ρ **and** source location/mechanism

surface wave ray coverage



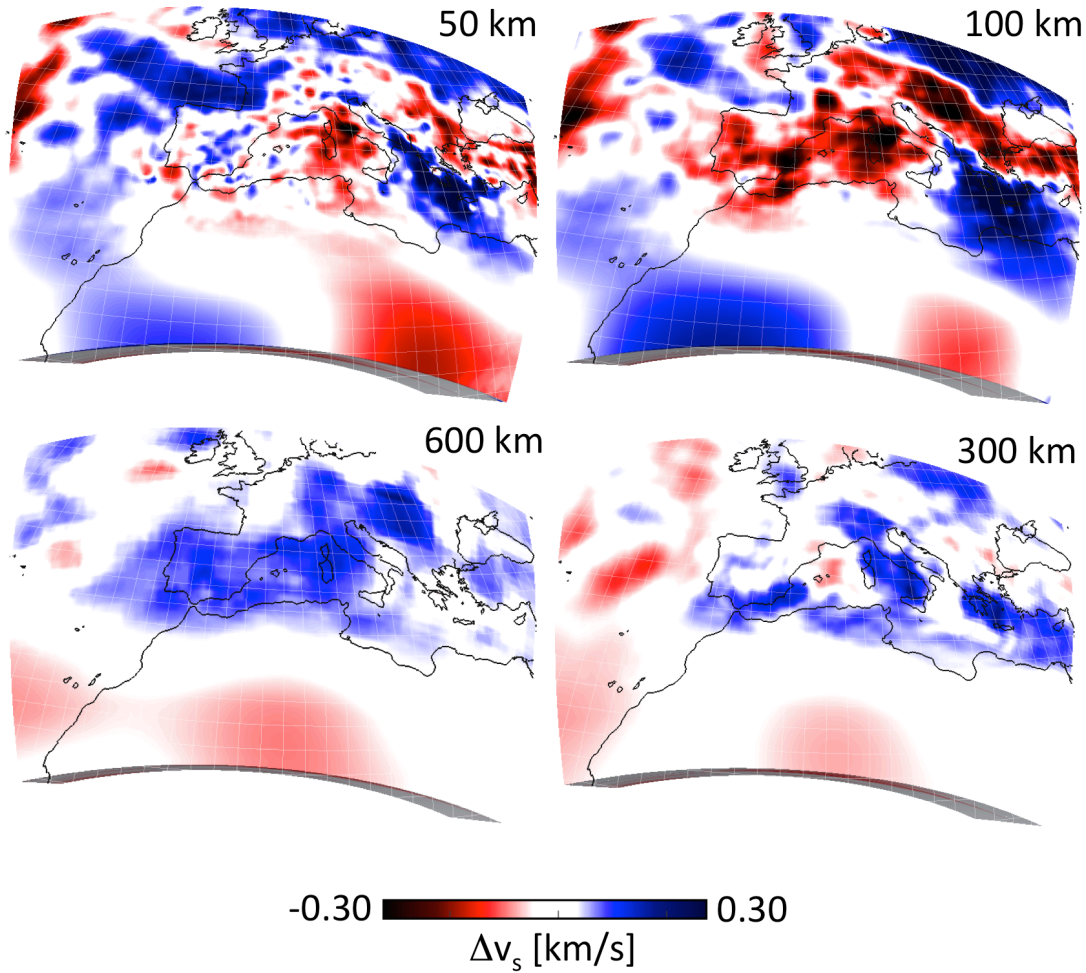
S VELOCITY MODEL

isotropic S velocity

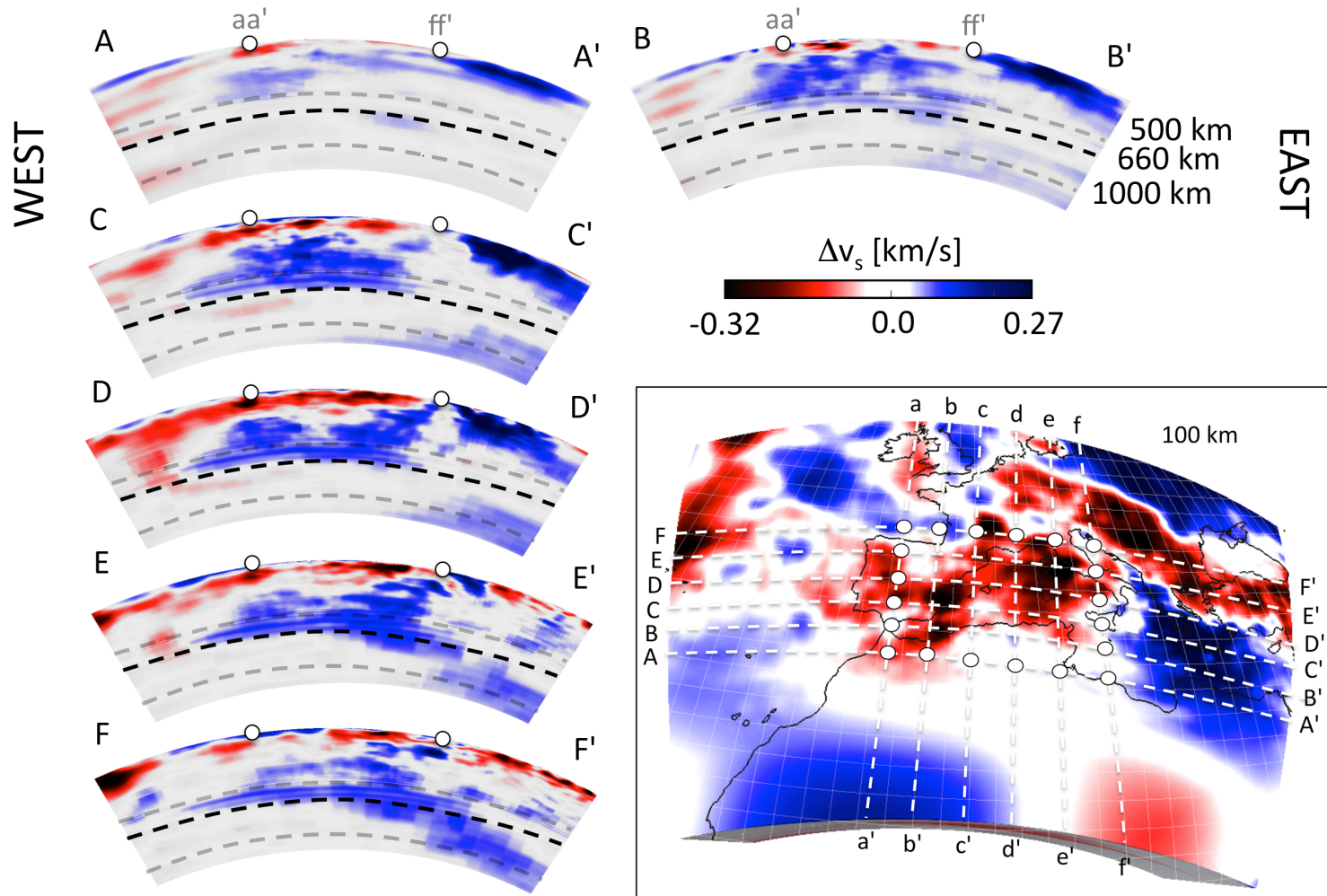


S VELOCITY MODEL

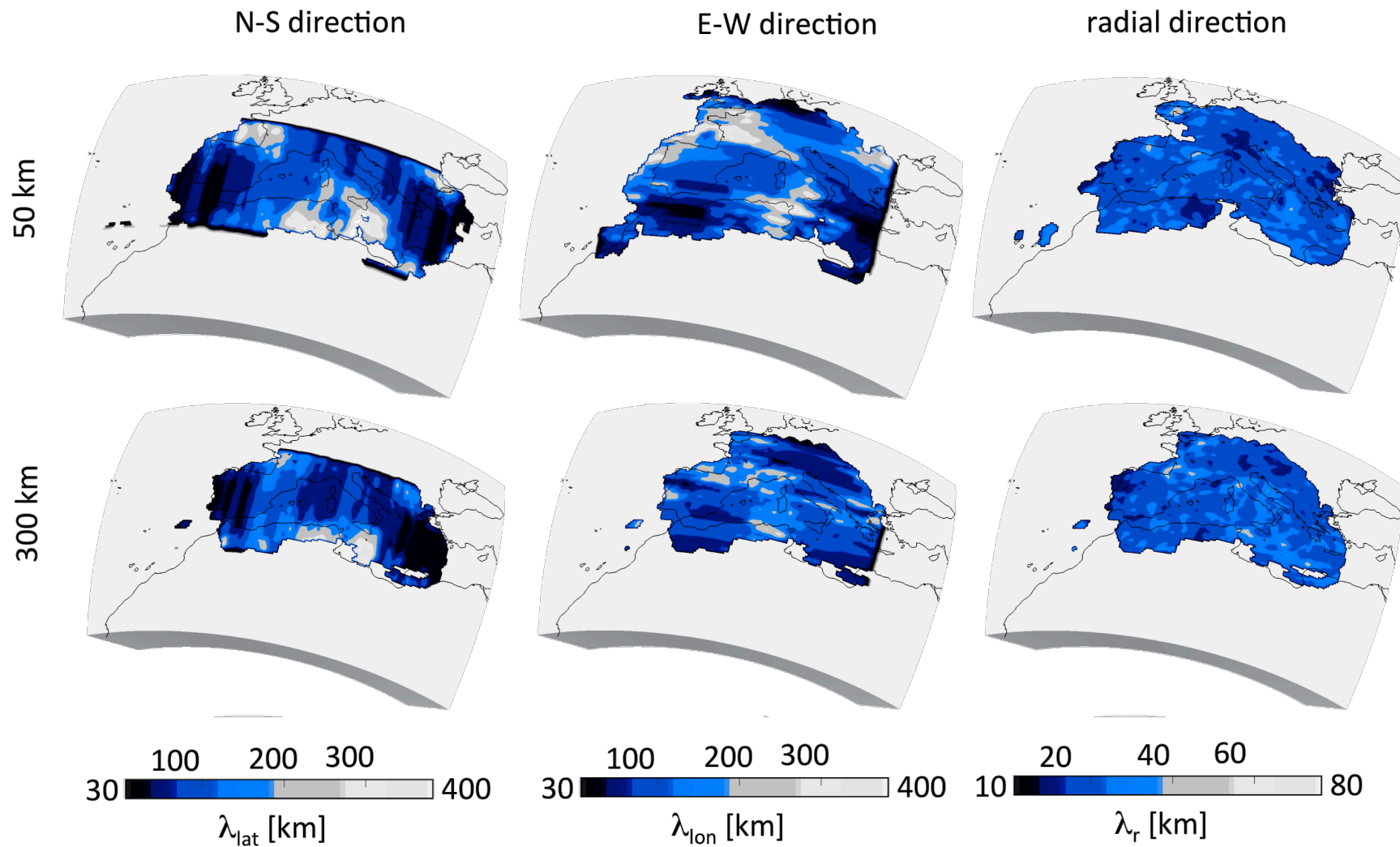
isotropic S velocity variations



S VELOCITY MODEL



POSITION- AND DIRECTION-DEPENDENT RESOLUTION LENGTHS



Conclusions

Limitations:

1. Local analysis

Limitations:

1. Local analysis

Benefits:

1. Quantify spatial resolution and inter-parameter trade-offs.

Limitations:

1. Local analysis

Benefits:

1. Quantify spatial resolution and inter-parameter trade-offs.
2. Low computational costs
 - around 5 Hessian-model applications
 - equivalent to around 5 CG iterations
 - much less than a synthetic inversion

Limitations:

1. Local analysis

Benefits:

1. Quantify spatial resolution and inter-parameter trade-offs.
2. Low computational costs
 - around 5 Hessian-model applications
 - equivalent to around 5 CG iterations
 - much less than a synthetic inversion
3. Low algorithmic complexity
 - easy to implement without modifications of existing codes

Limitations:

1. Local analysis

Benefits:

1. Quantify spatial resolution and inter-parameter trade-offs.
2. Low computational costs
 - around 5 Hessian-model applications
 - equivalent to around 5 CG iterations
 - much less than a synthetic inversion
3. Low algorithmic complexity
 - easy to implement without modifications of existing codes
4. Scalability
 - 5 random models sufficient in 1, 2 and 3 dimensions [empirical]

Limitations:

1. Local analysis

Benefits:

1. Quantify spatial resolution and inter-parameter trade-offs.
2. Low computational costs
 - around 5 Hessian-model applications
 - equivalent to around 5 CG iterations
 - much less than a synthetic inversion
3. Low algorithmic complexity
 - easy to implement without modifications of existing codes
4. Scalability
 - 5 random models sufficient in 1, 2 and 3 dimensions [empirical]
5. **Applicability to any tomographic technique**

Thanks for your attention!